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### **Detection and Classification from** Hyperspectral Imagery Using the Normal Compositional Model

David Stein

**ASAP 2003** 

11-13 March 2003

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#### **Outline**

- Hyperspectral Imaging (HSI) aka Imaging Spectrometry
- Descriptive models of HSI
- The Normal Compositional Model
- **Applications**
- Summary
- Future Work



### Hyperspectral Imaging or Imaging Spectrometry

NASA: ER2 AVIRIS



**NASA: E01 HYPERION** 



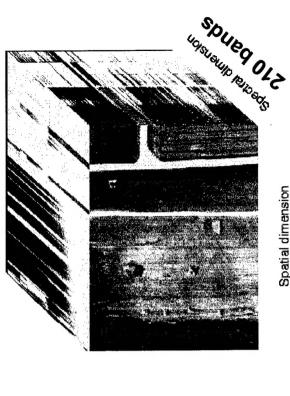
Hyperspectral Imager



NRL: HYDICE 0.4-2.4 nm

Dispersing element

Entrance slit



Detector array

Spatial axis

Sensor FOV

Scan direction (platformention or mirror)

Spatial dimension

308 pixels

əonofoəlfər Q 4

reflectance spectra of

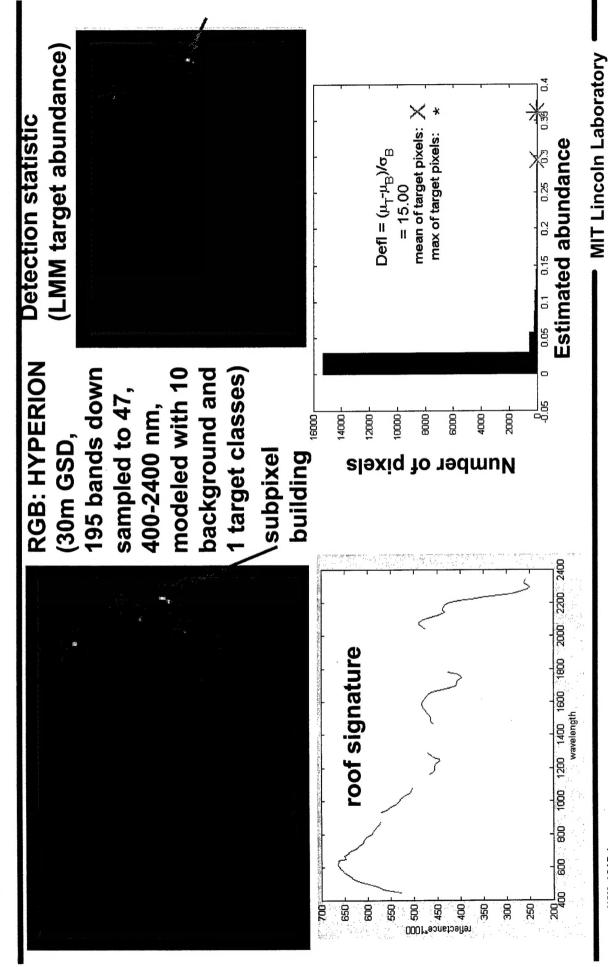
various soils

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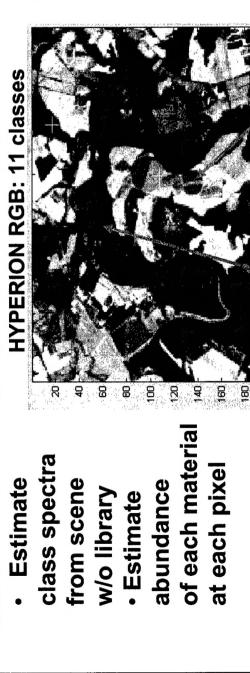
## **Detection of a Known Target**

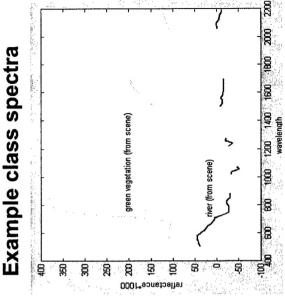


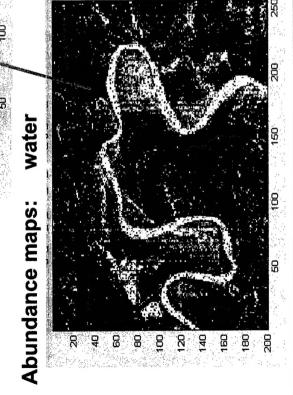
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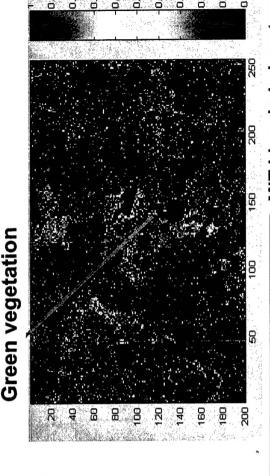


## **HSI: Blind Unmixing**









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#### Outline

Hyperspectral Imaging (HSI) aka Imaging Spectrometry

Descriptive models of HSI

Characteristics of HSI

Modeling intra-class variability

Common approaches to modeling HSI

**Applications** 

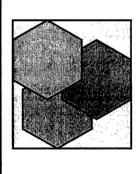
Summary

Conclusions



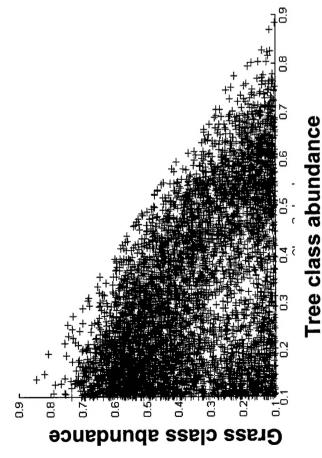
## Important Characteristics of HSI

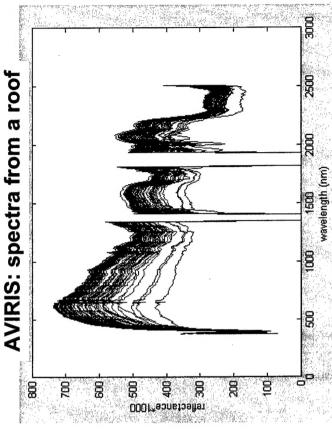
Different materials occlude each other



material 1material 2material 3background

Pixels are generally mixtures (ARMY Night Vision Lab: NVIS- forest)



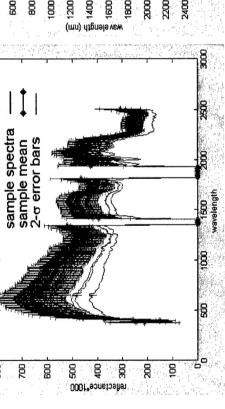


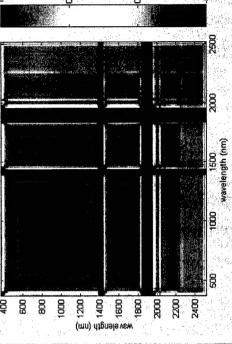
Materials not well modeled by a single spectrum

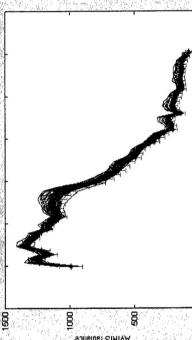


## Random Models of Spectral Variability: First and Second Order Statistics

AVIRIS
(400-2400 nm)
Southern CA
multiple roof
reflectance







radiance spectra

(400-900 nm)

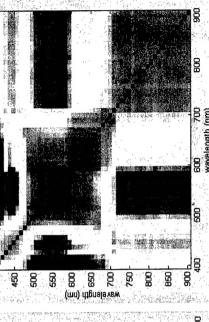
**AVIRIS** 

Tampa Bay

identified with

phytoplankton

class



- Correlation matrix is independent of radiance-to-reflectance transformation
- Correlation matrix is class dependent
- MIT Lincoln Laboratory Significance of modeling variability judged by impact on performance

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# Subspace Models of Spectral Variability

- Subspace model:
- Define a low-dimensional subspace such that target signatures may be replaced, with bounded error, by projection onto subspace
- Eigenvector construction

Given observations 
$$\{x_i, \dots, x_m\} \subset R^n$$
 define  $T = \sum_{i=1}^m x_i \cdot x_i^* = UDU^n$ 

Relative magnitude of the error vector obtained by ignoring the last N-p eigenvectors, where N=rank(T), is



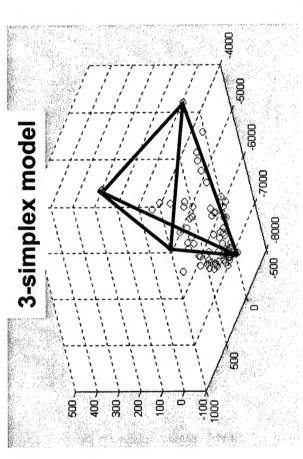
Applied to the roof data above the roof data a

$err^2(p)$	60000	0.0001
Dimension (p)	9	0



# Convex Models of Spectral Variability

- as a convex mixture (linear combination such that coefficients are Find simplex such that every target is approximately represented positive and sum to 1) of the vertices
- Construction of maximum volume inscribed n-simplex
- Project samples onto first n eigenvectors of correlation matrix
- Select n+1 affine independent samples
- apply determinant update equations to maximize volume



roof data

**AVIRIS** 

\* Endmember

Projected data



# Common Approaches to Modeling HSI

Local Normal model

$$x \in Neigh(z) \Rightarrow x \sim N(\mu_z, \Gamma_z)$$

Normal mixture model

$$x \sim \sum_{j=1}^{m} \rho_j N(\mu_j, \Gamma_j), \rho_j \geq 0 \text{ and } \sum_{j=1}^{m} \rho_j = 1$$

Subspace models (linear)

$$x_i = A\alpha_i + n, n \sim N(\mu, \Gamma)$$

Linear mixture models (convex)

$$x_i = \sum_{j=1}^{m} a_{ij} s_j + n; a_{ij} \ge 0 \text{ and } \sum_{j=1}^{m} a_{ij} = 1; n \sim N(\mu, \Gamma)$$

 None of these models accounts for 1) occlusion, 2) intra-class variability, and 3) subpixel mixing



#### Outline

- Hyperspectral Imaging (HSI) aka Imaging Spectrometry
- Descriptive models of HSI
- The Normal Compositional Model
- Applications
- Classification
- Detection
- Summary
- Future Work



## Normal Compositional Model

• Observation  $\vec{x}_i$  is modeled as

$$\vec{x}_i = \vec{e}_0 + \sum_{j=1}^m a_{ij} \vec{e}_j$$
 such that  $\vec{e}_0 \sim N(\vec{u}_0, \Gamma_0)$  and  $\vec{e}_j \sim N(\vec{u}_j, \Gamma_j)$ 

subject to constraints

$$0 \le a_{ij}$$
 for  $j \le r$ , and either  $\sum_{j=1}^{r} a_{ij} = 1$  or  $\sum_{j \ne 1}^{r} a_{ij} \le 1$ ;  $r \le m$ .

Features:

Estimates of class parameters converge under appropriate hypotheses to Models subpixel mixing and random variation within a class Class parameters  $\{(\mu_j,\Gamma_j)|0 \le j \le m\}$  estimated as scene wide aggregates

Abundance values  $\{a_{ij} | 1 \le j \le m, 1 \le i \le N\}$  estimated at every pixel

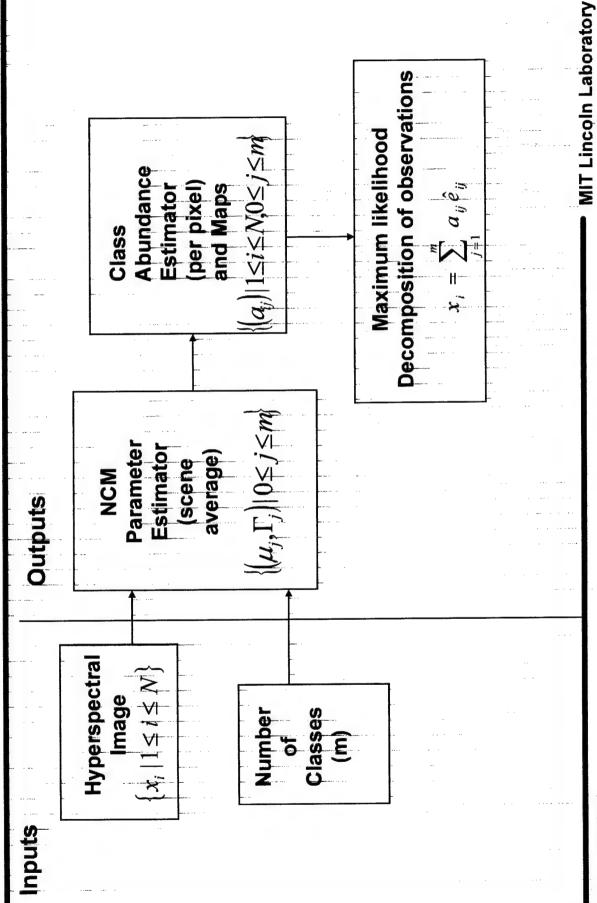
Additive components, e.g. noise and path radiance

Accommodates (fat) subspace as well as (fat) simplex models.

Special Cases: Linear mixture, Gaussian mixture and subspace models

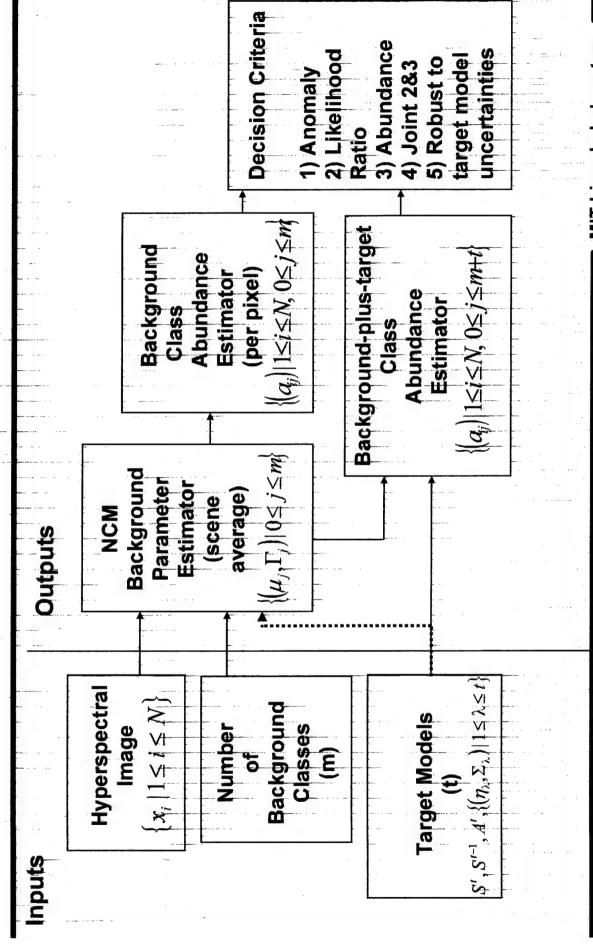


## **NCM Blind Unmixing**





### **NCM Detection**



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## **Nested Expectation Maximization Estimation of NCM Parameters:**

- Complete Likelihood function:
- N observations x, and abundance vectors  $\alpha_i = (\alpha_{i1}, ..., \alpha_{im})$

$$p(x_1, \dots, x_N, a_{11}, \dots a_{1m}, \dots, a_{N1}, \dots, a_{Nm} \mid \{(\mu_j, \Gamma_j)\}\} = \prod_{i=1}^N N(x_i; \mu(\alpha_i) + \mu_0, \Gamma(\alpha_i) + \Gamma_0) p(\alpha_i)$$

where

$$\mu(\alpha_i) = \sum_{ij} a_{ij} \mu_j$$
 and  $\Gamma(\alpha_i) = \sum_{ij} a_{ij}^2 \Gamma_j$ 

- Abundance values are hidden parameters  $\{(\mu^{\circ}, \Gamma^{\circ})|_{0 \le j \le m}\}$
- Linear mixture model techniques to identify initial endmembers (HSI) Vertices of convex hull (low dimensional, e.g., multispectral, data)
  - 1. Sample hidden parameters ⟨a, ′ |0≤ j ≤ m, 1≤ i ≤ N
- Optimization of likelihood function (currently)
  - Monte Carlo Markov Chain (in progress)
- 2. Optimize class parameters for given values of hidden parameters Expectation-Maximization Algorithm  $\{(\mu_j)^{k-1} | \Gamma_j^{\lambda_j} \} | 0 \le j \le m \}$
- 3. Repeat 1 and 2 until a convergence criterion is met



## Updating Class Parameters Using **Expectation Maximization**

Class parameters after iteration \( \ell \)

$$\Omega^{\ell} = \left\{ \left( \mu_j^{\ell}, \Gamma_j^{\ell} \right) | \ 0 \le j \le m \right\}$$

Class mean update: 
$$\mu_{k+1} = \frac{1}{N} \sum_{i=1}^{N} E(e_k \mid x_i, \{a_{ij}^r\}, \Omega^s)$$

$$= \mu_k + \frac{1}{N} \sum_{i=1}^{N} a_{ii} [\Gamma_k'[\Gamma'(\alpha_i) + \Gamma_0'] (x_i - \mu'(\alpha_i) - \mu_0')$$

Class covariance update

$$\Gamma_k^{\prime + 1} = \frac{1}{N} \sum_{i=1}^N {\sf cov}(e_k \mid x_i, \Omega^{\prime \prime}) + \left[ E(e_k \mid x_i, \Omega^{\prime \prime}) - \mu_k^{\prime + 1} \right] E(e_k \mid x_i, \Omega^{\prime \prime}) - \mu_k^{\prime + 1} \right]$$

Parameter updates are averages over expected values that are calculable from current parameters and abundance values



#### Outline

- Hyperspectral Imaging (HSI) aka Imaging Spectrometry
- Descriptive models of HSI
- The Normal Compositional Model
- Generalization of common models
- Estimation
- Classification
- Detection
- **Applications**
- Summary
- **Future Work**



# AVIRIS Imagery of Cuprite, Nevada

## AVIRIS: RGB Cuprite, NV



Acid sulfate hydrothermal alteration centers created 7.6-6.2 million years ago (hot sulfuric acid laden water flowing through surrounding rock changes the mineral content)

Complex well studied scene used for evaluating algorithms

Mineral classification maps and spectral library available from US Geological Survey

USGS airborne hyperspectral identifications confirmed using ground spectrometry and laboratory analysis of field samples

19 minerals plus 4 mixtures identifiable using SWIR data over 189 km² area

Subtle shifts in signatures due to variations in constituent elements, crystalline structure, temperature of formation

Validate NCM estimation and blind unmixing

Sensor: AVIRIS

Spectral region: 50 bands (2-2.5µm)

Area covered: 42 km², 350 by 300 pixels extracted from a 189 km² image

Initial number of classes: 18



### Identification of Spectra: Matching Absorption Features

Alunite Group:

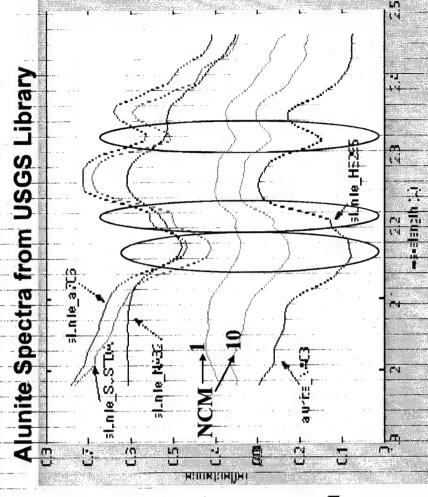
 $(Na, K)Al_3(SO_4)_2(OH)_6$ 

- **Absorption Features:**
- Increasing concentration of Na shifts shoulder longer, and main band narrows

Shoulder: O-H stretch+Al-O-H

oend

− 2.31μ : O-H stretch+Al-O-H bend



matching technique (correlation coefficients 0.99 and 0.98, respectively) NCM mean spectra identified with high temperature K-alunite (10) and moderate temperature mixed alunite (1) using USGS feature



#### USGS Class Map and NCM Alunite Abundance Plane

Good qualitative comparison/ feature identification supports NCM unmixing Alumet Zeolule JSGS Class Map K-Alm4e 1500 K-Adurate 4500 NCM abundance: sum of classes 1&10

11 identified with minerals, 3 mixtures, 4 Fe bearing minerals

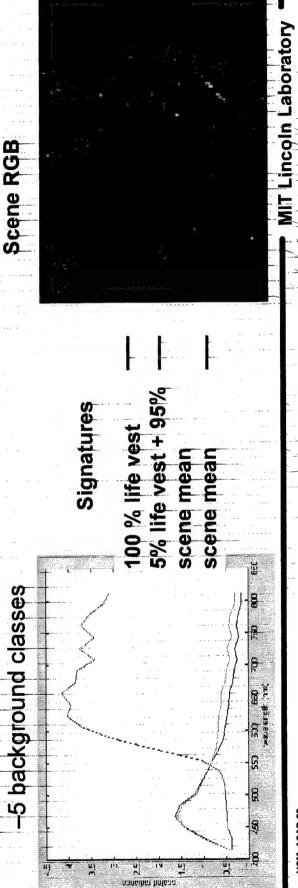
Of 18 classes,

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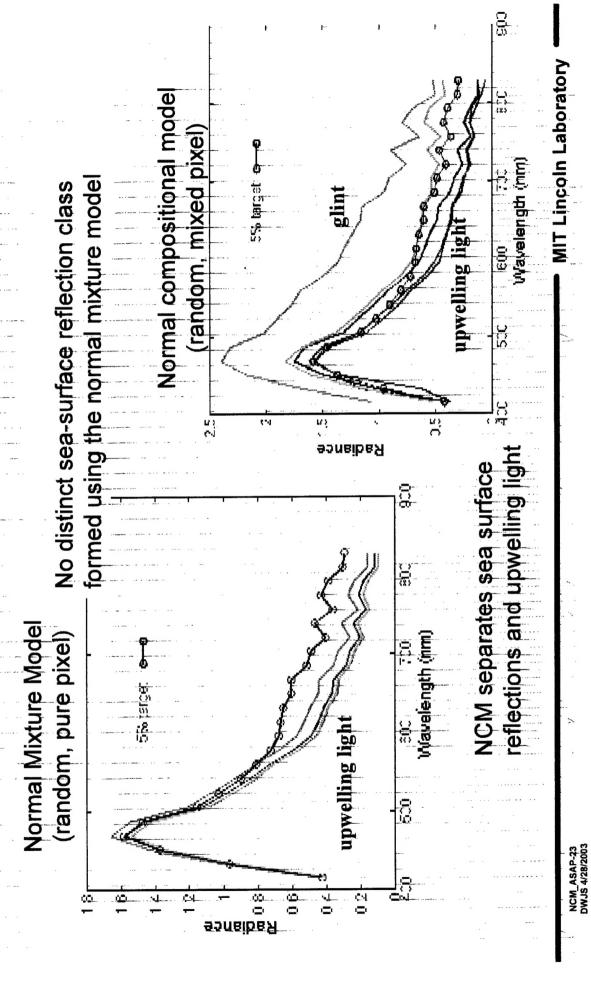
#### Life Vests in Ocean HSI **Detection Experiment:**

- Compare detection performance of Gaussian mixture, linear mixture, and NCM based known target and anomaly detection algorithms
- Background Data
- -125-by-125 2 m<sup>2</sup> pixels
- 24 band VNIR HSI (415-830 nm) from LASH sensor
  - Target description
- Life vest signature combined with background data from 1000 randomly selected pixels at 5% pixel fill fraction
  - –1 target class (mean given, covariance estimated as noise covariance)
    - Model Parameters
- -5 background classes



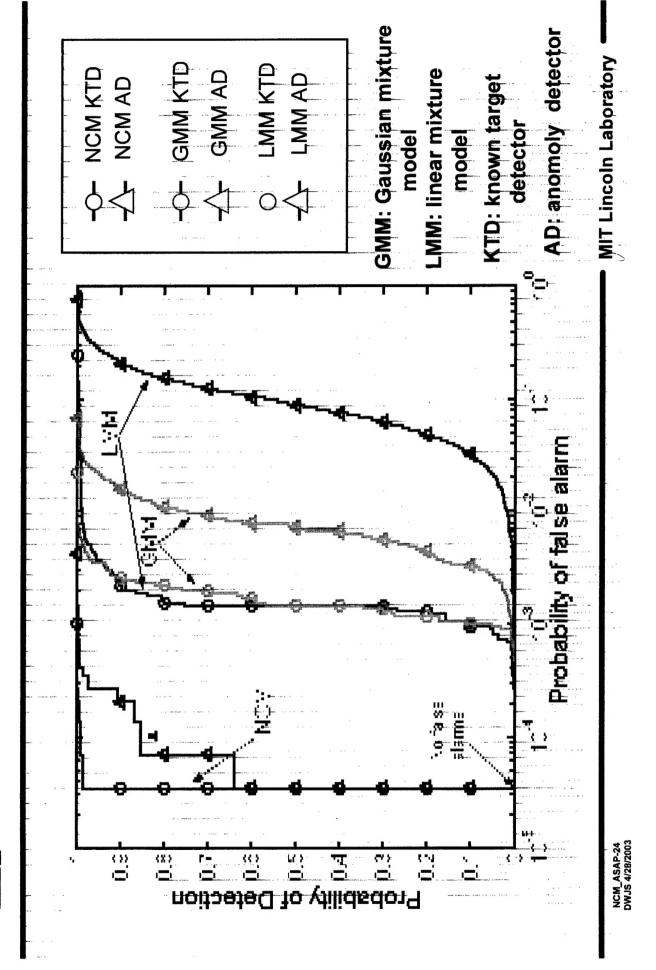
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# **Comparative Detection Performance**





## **Summary and Conclusions**

Described a normal compositional model (NCM)

simultaneously treats mixed pixels and intra-class variability

accommodates subspace, convex, and random class variability

generalizes and synthesizes normal mixture, linear mixture, and subspace models

Applied NCM to Cuprite data

Class means identified with spectra of materials in scene

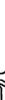
Abundance estimates qualitatively corresponded with USGS classification maps

Application to ocean data

NCM estimation method identified classes where pure-pixel methods failed

NCM offered superior detection performance in comparison with LMM and GMM based models

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#### **Future Work**

Speed up the software

Applications to real time HSI systems

Applications to HSI performance models

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